

PHOTOELECTRIC PHOTOMETRY OF THE ECLIPSING BINARY V 505 MON

D. Chochol

Astronomical Institute, Slovak Academy of Sciences
059 60 Tatranská Lomnica, Czechoslovakia

G. A. Bakos

Department of Physics, University of Waterloo
Waterloo, Ontario, Canada N2L 3G1

C. Bartolini, A. Guarnieri

Department of Astronomy, Bologna University
40100 Bologna, Italy

A. Dapergolas

Halkys 59, Nea Ionia
Athens, Greece

L. Szabados

Konkoly Observatory of the Hungarian Academy of Sciences
1525 Budapest XII, Hungary

Received 14 September 1984

ABSTRACT. Photoelectric U, B, V observations of the eclipsing binary V 505 Mon, performed at the observatories in Skalnaté Pleso, Budapest, Bologna and Waterloo in the years 1972-1984 are presented. The following ephemeris has been derived, using all these data:

$$J.D.^{\text{prim.min.}} = 2441328.06 + 53.7675^d E .$$

Indications of ongoing mass transfer in a semidetached binary configuration are

presented. The possible causes of the observed short-term changes of brightness are discussed.

ФОТОЭЛЕКТРИЧЕСКАЯ ФОТОМЕТРИЯ ЗАТМЕННОЙ ПЕРЕМЕННОЙ V 505 ЕДИНОРОГА. В работе опубликованы фотоэлектрические U, B, V наблюдения затменной переменной V 505 Единорога, полученные на обсерваториях Скалнате Плесо, Будапешт, Болонья и Ватерлоо во время 1972-1984 гг. Все эти наблюдения использованы для выведения следующей эфемериды:

$$J.D. \text{ ГЛ.МИН.} = 2441328,06 + 53,7675 E .$$

Приведены признаки переноса вещества в полуразделенной двойной системе. Дискутируются причины наблюдаемых коротковременных изменений блеска.

FOTOELEKTRICKÁ FOTOMETRIA ZÁKRYTOVEJ DVOJHVIEZDY V 505 MON. V práci sú publikované fotoelektrické U, B, V pozorovania zákrytovej dvojhviezdy V 505 Mon, získané v rokoch 1972-1984 na observatóriach Skalnaté Pleso, Budapešť, Bologna a Waterloo. Tieto údaje boli použité na odvodenie novej efemeridy:

$$J.D. \text{ prim.min.} = 2441328.06 + 53.7675^d E .$$

Pozorované náznaky prenosu hmoty v sústave vedú k domnieke, že ide o polodotykovú dvojhviezdu. V práci sú diskutované možné príčiny zistených krátkodobých zmien jasnosti.

1. INTRODUCTION

V 505 Mon (HD 48 914) is a relatively bright ($V = 7.23$, $B-V = 0.01$, $U-B = -0.43$) but until recently not very extensively studied Be star. Hoag and Smith (1959) classified the object as B5 Ib, Turner (1976) as B5 II, Chochol (1983) as B3 II-III. The light variations of V 505 Mon was discovered by Wachmann (1966). Chochol and Kučera (1981) proved, that V 505 Mon was an eclipsing binary with a period of 53.7805^d . The orbital period was confirmed spectroscopically by Stagni et al. (1982). According to them, V 505 Mon is a detached system with component masses of $m_1 = 50 M_\odot$ and $m_2 = 27 M_\odot$. De Gréve et al. (1983) found, however, that the spectral components were at variance with the other parameters in the HR diagram.

The purpose of this paper is to determine a new ephemeris of the eclipsing binary V 505 Mon, to discuss the asymmetries of the light curve and to search for the periodicity and cause of short-term changes in brightness, using the photometric material obtained at four different observatories during 1972-1984.

2. OBSERVATIONAL MATERIAL

The U, B, V photoelectric observations of V 505 Mon were performed with the 0.6 m telescopes of the observatories: Skalnaté Pleso, the Konkoly Observatory in Budapest, the University Observatory in Bologna and with 0.33 m te-

lescope of the University Observatory in Waterloo. The observations have been listed in Tab. 1, in which each entry represents the mean of 5-10 individual observations. The data were reduced to a standard system and are presented in terms of the differences V 505 Mon minus HD 48 956 ($V = 8.15$, $B-V = 0.03$, $U-B = -0.23$). On nights when only V observations were made, the data were reduced from the instrumental to the standard system using a statistically determined correction.

The program HEC 21 based on Stellingwerf's (1978) phase dispersion minimization method, given at our disposal by Dr. P. Harmanec from Ondřejov Observatory was used to determine the improved value of the period. The new ephemeris is

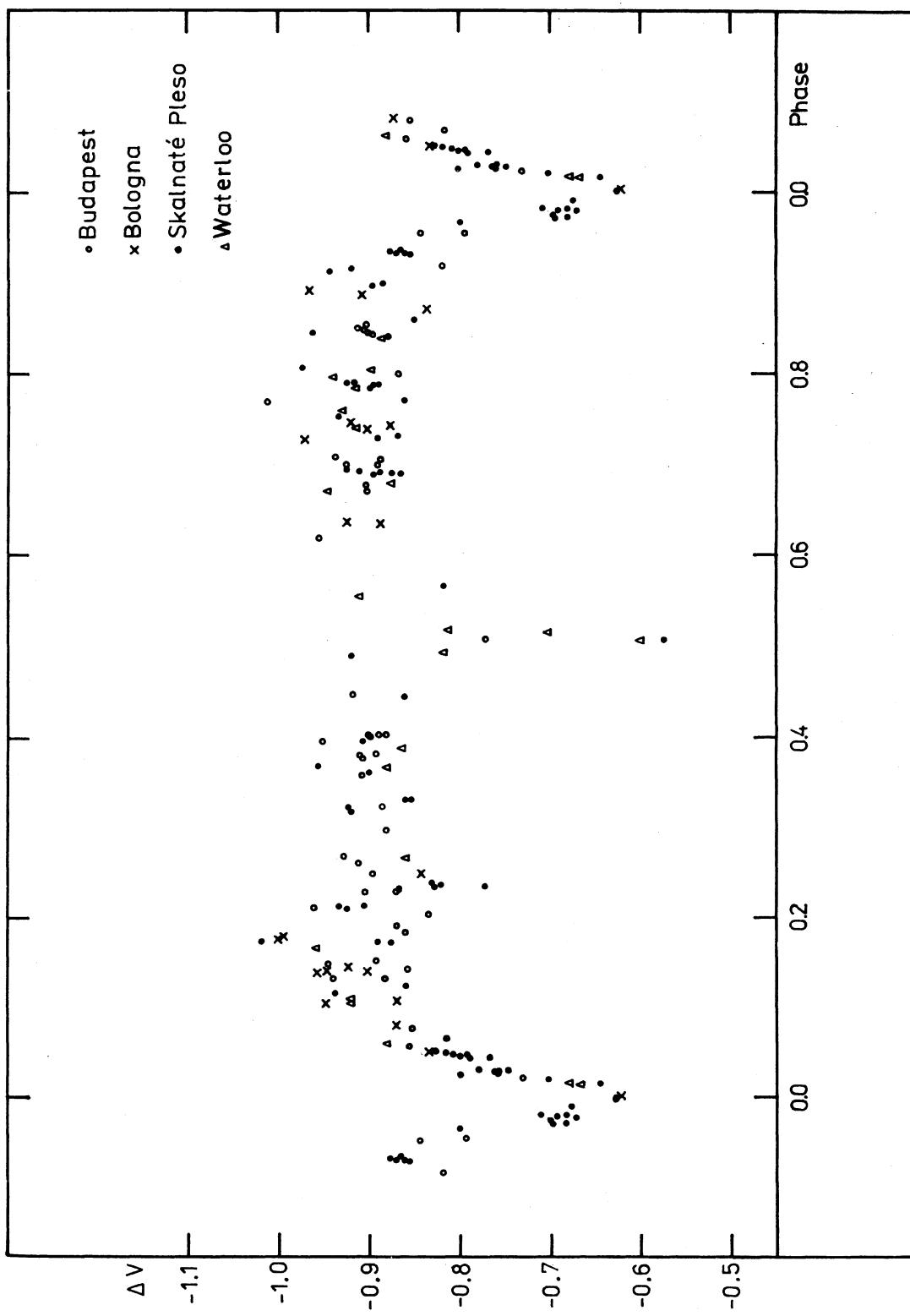
$$J.D. \text{ prim. min.} = 2441328.06 + 53.7675^d E . \quad (1)$$

The photoelectric observations given in Table 1 are averages over 0.05 day and plotted as a function of orbital phase in Fig. 1.

As can be seen from Fig. 1, the long period changes of brightness, expressed by ephemeris (1) are apparently related to the eclipses of the components. The light curve, however, does not resemble light curves of detached eclipsing binaries, because of the presence of asymmetries. The remarkable features are brightening after the primary minimum, the decrease of brightness before the primary minimum and the large scatter of observations in the secondary minimum. The light curve displays also short-term changes in brightness with an amplitude of 0.15^m . Consequently, a search for shorter periods has been undertaken after the removal of the 53.7675^d period. The program HEC 21 indicated two possible periods 0.757065^d and 1.581002^d . However, the phase plots of the residual data with these periods are not very convincing, leaving thus doubts about the presence of a real short period.

3. DISCUSSION

Short-term brightness variations are frequently observed in Be stars. In some cases the period of variations was found and it is generally accepted that variations are associated with rotational or pulsational period of the star in question (Harmanec, 1983 a,b). The amplitude of short-term variations in V 505 Mon is one of the largest observed in Be stars (the amplitudes of the brightness variations are mostly less than 0.1^m). The study of physical nature of variations may have a fundamental significance for understanding of Be phenomenon. The study of the periodicity of short-term brightness variations of V 505 Mon using the presented material may be affected by the fact, that changes may have occurred in the circumstellar envelope during the 12 years covered by observations. This could have an effect on the brightness of the object. The differences in the spectral classification of V 505 Mon indicate that this may be the case. It is interesting to note, that the brightening of the V magnitude of V 505 Mon out of the eclipses is accompanied by reddening of the B-V and blueing of the U-B indices. It indicates type 2 long term colour variations described by Harmanec (1983 a). In this case the object changes its photometric



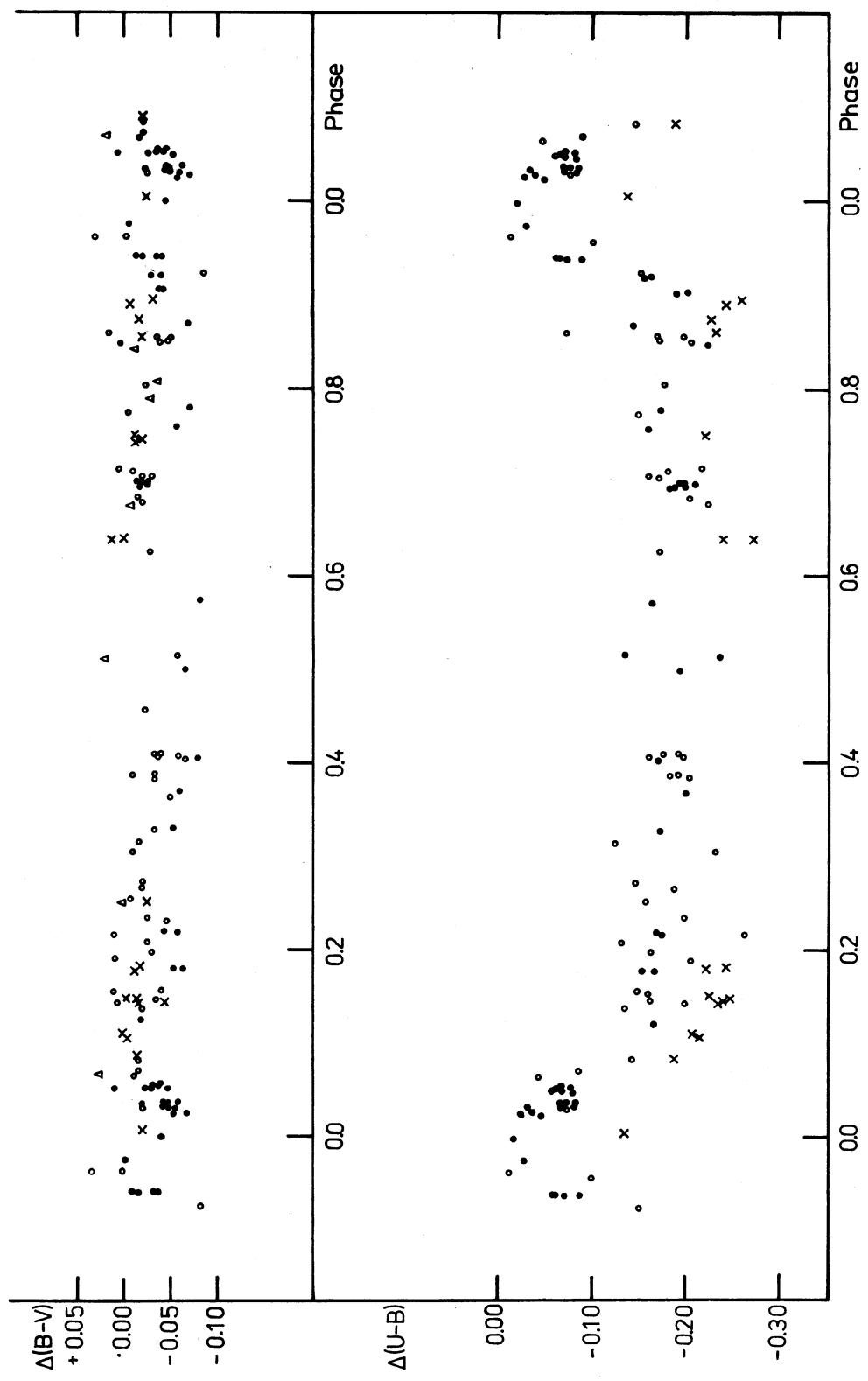


Fig. 1. U, B, V photometry of V 505 Mon.

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
Bologna									
2088.39	0.1411	-0.899	-0.051	-0.142	4278.45	0.8731	-0.830	-0.023	-0.227
.44	0.1420	-0.906	-0.052	-0.134	4279.27	0.8884	-0.907	-0.012	-0.243
2785.45	0.1054	-0.948	-0.009	-0.225	.45	0.8917	-0.948	-0.044	-0.260
.49	0.1062	-0.947	-0.014	-0.210	.46	0.8919	-0.991	-0.032	-0.266
2787.35	0.1407	-0.958	-0.020	-0.242	4285.41	0.0026	-0.623	-0.028	-0.138
.37	0.1411	-0.963	-0.020	-0.234	4534.56	0.6364	-0.905	0.000	-0.237
.40	0.1417	-0.955	-0.026	-0.241	.563	0.6365	-0.867	+0.013	-0.224
.48	0.1432	-0.960	-0.021	-0.242	.572	0.6366	-0.924	-0.006	-0.273
.55	0.1445	-0.939	-0.023	-0.242	4539.55	0.7292	-0.957		
.57	0.1448	-0.935	-0.018	-0.240	.58	0.7298	-0.983	+0.021	-0.309
.59	0.1452	-0.932	-0.021	-0.240	4540.56	0.7480	-0.911	-0.020	-0.222
.61	0.1456	-0.914	-0.032	-0.235	.57	0.7482	-0.928	-0.016	-0.219
.62	0.1458	-0.917	-0.025	-0.232	4558.49	0.0815	-0.861	-0.025	-0.188
2789.36	0.1781	-1.010	-0.018	-0.186	.50	0.0816	-0.880	-0.020	-0.191
.44	0.1796	-0.987	-0.029	-0.259	4567.52	0.2494	-0.852	-0.024	-0.232
.46	0.1800	-0.990	-0.031	-0.242	.53	0.2496	-0.826	-0.035	-0.209
.47	0.1802	-0.997	-0.024	-0.246	.54	0.2498	-0.852	-0.039	-0.242
.49	0.1805	-1.004	-0.017	-0.253					
2836.41	0.0532	-0.833			Budapest				
2839.38	0.1084	-0.871	-0.007	-0.210	1329.3851	0.0246	-0.731	-0.030	-0.073
4271.30	0.7401	-0.901	-0.020		.4309	0.0255	-0.731	-0.026	-0.081
.31	0.7403	-0.905	-0.019		1331.3493	0.0612	-0.871	-0.017	-0.030
.50	0.7439	-0.878	-0.024		.4583	0.0632	-0.844	-0.023	-0.075
4277.39	0.8534	-0.898	-0.024	-0.231	1332.3766	0.0803	-0.855	-0.024	-0.136
.41	0.8538	-0.909	-0.025	-0.234	.4989	0.0826	-0.854	-0.023	-0.155
4278.42	0.8726	-0.838	-0.015	-0.226	1340.3441	0.2285	-0.905	-0.053	
.44	0.8729	-0.839	-0.028	-0.211	1352.3411	0.4516	-0.917	-0.029	

Table 1 (continued)

J.D. hel 244000+	Phase	ΔV	$\Delta (B-V)$	$\Delta (U-B)$	J.D. hel 244000+	Phase	ΔV	$\Delta (B-V)$	$\Delta (U-B)$
1364.3722	0.6754	-0.902	-0.013	-0.235	1617.6112	0.3852	-0.874	-0.053	-0.186
.4149	0.6762	-0.903	-0.037	-0.184	1672.3923	0.4041	-0.868	-0.084	-0.169
1366.3377	0.7119	-0.935	0.000	-0.217	*3986	0.4042	-0.872	-0.095	-0.159
1389.3552	0.1400	-0.940	-0.001	-0.202	*4048	0.4043	-0.880	-0.090	-0.161
1392.3074	0.1949	-0.870	-0.038	-0.166	*4118	0.4045	-0.914	-0.046	-0.189
1393.3131	0.2136	-0.961	+0.004	-0.265	*4180	0.4046	-0.923	-0.043	-0.125
1394.3079	0.2321	-0.870	-0.032	-0.201	*4236	0.4047	-0.909	-0.077	-0.184
1395.3206	0.2510	-0.896	-0.014	-0.160	*4312	0.4048	-0.905	-0.049	-0.125
1396.3702	0.2705	-0.928	-0.026	-0.148	*4375	0.4049	-0.928	-0.030	-0.199
1399.3193	0.3253	-0.885	-0.040		*4437	0.4051	-0.919	-0.054	-0.186
1401.3124	0.3624	-0.909	-0.056		*4500	0.4052	-0.944	-0.016	-0.199
1402.3238	0.3812	-0.908	-0.041	-0.205	*4576	0.4053	-0.935	-0.025	-0.220
1403.3399	0.4001	-0.950	-0.073		*4645	0.4054	-0.853	-0.053	-0.230
1415.3290	0.6231	-0.955	-0.032	-0.173	*4858	0.4058	-0.871	-0.059	-0.193
1594.6084	0.9574	-0.885	-0.010	-0.083	*4913	0.4059	-0.896	-0.044	-0.170
.6147	0.9575	-0.850	-0.006	-0.114	*4979	0.4061	-0.901	-0.037	-0.207
.6223	0.9577	-0.794	-0.060	-0.099	*5027	0.4062	-0.890	-0.051	-0.229
.6278	0.9578	-0.844	-0.010	-0.099	*5083	0.4063	-0.905	-0.042	-0.216
.6322	0.9579	-0.826	-0.028	-0.117	*5135	0.4064	-0.906	-0.031	-0.212
1617.5340	0.3838	-0.930	0.000	-0.216	*5194	0.4065	-0.888	-0.043	-0.217
.5418	0.3840	-0.919	-0.017	-0.184	*5249	0.4066	-0.886	-0.052	-0.190
.5480	0.3841	-0.912	-0.030	-0.187	*5305	0.4067	-0.875	-0.052	-0.189
.5557	0.3842	-0.910	-0.011	-0.201	*5375	0.4068	-0.867	-0.059	-0.179
.5633	0.3838	-0.886	-0.023	-0.172	*5430	0.4069	-0.830	-0.086	-0.162
.5709	0.3845	-0.899	-0.023	-0.187	*5479	0.4070	-0.896	-0.021	-0.160
.5775	0.3846	-0.912	-0.021	-0.193	*5538	0.4071	-0.913	-0.021	-0.180
.5848	0.3848	-0.899	-0.036	-0.196	*5597	0.4072	-0.884	-0.035	-0.194
.6050	0.3851	-0.877	-0.065	-0.162	*5652	0.4073	-0.896	-0.013	-0.227

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 244000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
1672.5705	0.4074	-0.869	-0.059	-0.230	1742.2701	0.7037	-0.967	+0.029	-0.143
.5756	0.4075	-0.880	-0.059	-0.211	.2737	0.7038	-0.935	+0.002	-0.156
1696.3085	0.8489	-0.917	-0.024	-0.209	.2793	0.7039	-0.904	-0.038	-0.184
.3143	0.8490	-0.895	-0.058	-0.243	.2835	0.7040	-0.964	+0.002	-0.216
.3206	0.8491	-0.879	-0.060	-0.206	.2876	0.7041	-0.970	+0.023	-0.212
.3265	0.8492	-0.923	-0.027	-0.208	.2911	0.7041	-0.954	+0.010	-0.207
.3317	0.8493	-0.894	-0.033	-0.205	.2953	0.7042	-0.903	-0.040	-0.166
.3372	0.8494	-0.892	-0.045	-0.172	.2987	0.7043	-0.914	-0.050	-0.131
.3432	0.8496	-0.895	-0.053	-0.167	.3036	0.7043	-0.886	-0.048	-0.184
.3508	0.8497	-0.921	-0.028	-0.164	.3071	0.7044	-0.892	-0.033	-0.261
.3550	0.8498	-0.917	-0.062	-0.151	.3119	0.7045	-0.908	-0.021	-0.172
.3613	0.8499	-0.905	-0.056	-0.166	.3168	0.7046	-0.925	-0.017	-0.151
.3682	0.8500	-0.879	-0.054	-0.196	.3210	0.7047	-0.897	-0.049	-0.140
.3731	0.8501	-0.898	-0.059	-0.195	.3244	0.7047	-0.869	-0.021	-0.192
.4053	0.8507	-0.905	-0.029	-0.187	.3286	0.7048	-0.867	-0.040	-0.165
.4099	0.8508	-0.907	-0.041	-0.207	.3314	0.7049	-0.856	-0.049	-0.165
.4143	0.8509	-0.913	-0.031	-0.209	.3369	0.7050	-0.910	-0.038	-0.133
.4190	0.8510	-0.889	-0.044	-0.213	1957.6153	0.7088	-0.908	-0.032	-0.164
.4251	0.8511	-0.904	-0.053	-0.186	.6197	0.7089	-0.886	-0.021	-0.205
.4310	0.8512	-0.907	-0.054	-0.167	.6239	0.7090	-0.884	-0.032	-0.177
.4357	0.8513	-0.928	-0.040	-0.171	.6278	0.7091	-0.887	-0.014	-0.195
.4411	0.8514	-0.904	-0.066	-0.164	.6316	0.7091	-0.877	-0.024	-0.205
.4455	0.8515	-0.904	-0.056	-0.171	.6361	0.7092	-0.871	-0.041	-0.173
.4509	0.8516	-0.921	-0.041	-0.171	.6410	0.7093	-0.893	-0.011	-0.160
.4564	0.8517	-0.912	-0.060	-0.183	1962.6328	0.8022	-0.867	-0.027	-0.178
1742.2557	0.7035	-0.829	-0.139	-0.056	1965.5823	0.8570	-0.937	+0.067	-0.054
.2612	0.7036	-0.861	-0.078	-0.131	.5873	0.8571	-0.961	+0.016	-0.006
.2647	0.7036	-0.976	+0.015	-0.171	.5922	0.8572	-0.899	-0.010	-0.083

Table 1 (continued)
 J.D. hel
 244000+ Phase ΔV $\Delta(B-V)$ $\Delta(U-B)$ J.D. hel
 244000+ Phase ΔV $\Delta(B-V)$ $\Delta(U-B)$

1965.5977	0.8573	-0.829	-0.029	-0.187	3219.3851	0.1760	-1.008
•6019	0.8574	-0.884	-0.032	-0.106	•3931	0.1761	-1.015
1980.5504	0.1354	-0.882	-0.031	-0.121	•4044	0.1764	-1.016
•5546	0.1355	-0.871	-0.035	-0.119	3436.6139	0.2161	-0.908
•5587	0.1356	-0.895	-0.011	-0.174	•6288	0.2164	-0.917
1981.5825	0.1546	-0.890	-0.030	-0.167	•6383	0.2166	-0.898
•5873	0.1547	-0.895	-0.067	-0.136	3437.6404	0.2352	-0.869
2108.3170	0.5117	-0.768	-0.066	-0.147	•6508	0.2354	-0.864
•3219	0.5118	-0.774	-0.063	-0.128	3464.5152	0.7351	-0.882
2465.3280	0.1516	-0.945	+0.003	-0.163	•5205	0.7352	-0.872
2467.2577	0.1875	-0.862	-0.001	-0.179	•5246	0.7352	-0.853
•2925	0.1881	-0.860	-0.003	-0.237	3467.4418	0.7895	-0.898
2468.2679	0.2063	-0.834	-0.032	-0.134	•4487	0.7896	-0.894
2471.3509	0.2636	-0.913	-0.028	-0.191	•4532	0.7897	-0.913
2473.4016	0.3017	-0.880	-0.015	-0.233	•4577	0.7898	-0.903
2767.5100	0.7717	-1.010	-0.008	-0.149	•4626	0.7899	-0.906
2816.3891	0.6808	-0.903	-0.019	-0.205	•4672	0.7900	-0.911
2829.3214	0.9214	-0.818	-0.089	-0.152	•4720	0.7901	-0.909
2831.3417	0.9589	-0.793	-0.006	-0.014	•4774	0.7902	-0.894
2837.2611	0.0690	-0.837	-0.004	-0.060	•4831	0.7903	-0.895
•3212	0.0701	-0.792	-0.041	-0.118	•4887	0.7904	-0.890
2841.3571	0.1452	-0.859	-0.042	-0.166	•4935	0.7905	-0.886
2850.3117	0.3117	-0.922	-0.022	-0.128	•4983	0.7905	-0.888
					•5046	0.7907	-0.904
					•5095	0.7908	-0.891
3219.3527	0.1754	-1.053			•5150	0.7909	-0.892
•3607	0.1755	-1.016			•5198	0.7909	-0.899
•3685	0.1757	-1.028			•5265	0.7911	-0.890
•3768	0.1758	-1.013			•5317	0.7912	-0.882

Skalnaté Pleso

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	J.D. hel 244000+	Phase	ΔV	J.D. hel 244000+	Phase	ΔV
3467.5373	0.7913	-0.886	3468.5623	0.8103	-0.980	3545.4178	0.2397	-0.823
.5428	0.7914	-0.892	.6059	0.8111	-0.972	.4269	0.2399	-0.822
.5478	0.7915	-0.900	.6115	0.8113	-0.971	.4839	0.2410	-0.830
.5525	0.7916	-0.901	.6185	0.8114	-0.958	.4941	0.2412	-0.832
.5588	0.7917	-0.895	3470.6743	0.8496	-0.967	4249.4605	0.3339	-0.864
.5637	0.7918	-0.895	.6789	0.8497	-0.961	.4705	0.3341	-0.856
.5685	0.7919	-0.899	3477.4193	0.9751	-0.701	.4787	0.3343	-0.861
.5738	0.7920	-0.901	.4281	0.9752	-0.695	.4880	0.3345	-0.863
.5780	0.7920	-0.895	.4651	0.9759	-0.686	.4971	0.3346	-0.859
.5831	0.7921	-0.894	.4742	0.9761	-0.697	.5054	0.3348	-0.855
.5887	0.7922	-0.897	.4894	0.9764	-0.665	.5140	0.3349	-0.854
.6014	0.7925	-0.896	.5003	0.9766	-0.676	.5234	0.3351	-0.854
.6063	0.7926	-0.900	.5093	0.9767	-0.683	4251.4439	0.3708	-0.947
.6122	0.7927	-0.914	.5175	0.9769	-0.691	.4523	0.3710	-0.946
.6171	0.7928	-0.924	.5266	0.9771	-0.688	.4606	0.3711	-0.958
.6223	0.7929	-0.928	.5566	0.9776	-0.700	.4689	0.3713	-0.966
.6275	0.7930	-0.929	.5713	0.9779	-0.691	.4787	0.3715	-0.962
.6323	0.7930	-0.925	3485.6009	0.1272	-0.857	.4883	0.3717	-0.959
.6375	0.7931	-0.912	.6105	0.1274	-0.862	.4968	0.3718	-0.965
.6420	0.7932	-0.912	3545.2513	0.2366	-0.777	4255.5771	0.4477	-0.862
.6469	0.7933	-0.917	.2596	0.2368	-0.769	4284.3414	0.9827	-0.693
.6517	0.7934	-0.923	.2951	0.2375	-0.831	.3497	0.9828	-0.681
.6566	0.7935	-0.927	.3033	0.2376	-0.818	.3581	0.9830	-0.680
.6614	0.7936	-0.927	.3120	0.2378	-0.806	.3665	0.9831	-0.685
.6685	0.7937	-0.932	.3215	0.2380	-0.852	.3750	0.9833	-0.702
.6761	0.7939	-0.936	.3298	0.2381	-0.836	.3858	0.9835	-0.710
.6816	0.7940	-0.923	.3900	0.2392	-0.824	.3935	0.9837	-0.673
3468.5511	0.8101	-0.985	.4004	0.2394	-0.824	.4031	0.9838	-0.677
.5567	0.8102	-0.981	.4094	0.2396	-0.824	.4123	0.9840	-0.701

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
4284.4233	0.9842	-0.681			4591.3933	0.6934	-0.890	-0.010	-0.198
4317	0.9844	-0.672			4026	0.6936	-0.907	-0.009	-0.185
4400	0.9845	-0.664			4100	0.6937	-0.896	-0.028	-0.200
4474	0.9847	-0.680			4158	0.6938	-0.891	-0.042	-0.198
4553	0.9848	-0.694			4222	0.6940	-0.890	-0.041	-0.195
4664	0.9850	-0.721			4288	0.6941	-0.888	-0.026	-0.207
4781	0.9852	-0.712			4358	0.6942	-0.899	-0.011	-0.200
4887	0.9854	-0.702			4433	0.6943	-0.893	-0.002	-0.224
4969	0.9856	-0.714			4500	0.6945	-0.889	-0.008	-0.235
4285.3745	0.0019	-0.636			4562	0.6946	-0.878	-0.027	-0.196
3828	0.0021	-0.632			4625	0.6947	-0.892	-0.010	-0.192
3913	0.0022	-0.626			4704	0.6949	-0.909	-0.007	-0.194
4001	0.0024	-0.622			4764	0.6950	-0.899	-0.044	-0.188
4085	0.0025	-0.623			4825	0.6951	-0.889	-0.045	-0.195
4168	0.0027	-0.616			4907	0.6952	-0.896	-0.018	-0.192
4251	0.0028	-0.626			4988	0.6954	-0.889	-0.041	-0.192
4552.5142	0.9703	-0.806	0.000	-0.021	5070	0.6955	-0.883	-0.056	-0.194
5273	0.9706	-0.788	-0.028	-0.040	5140	0.6957	-0.898	-0.012	-0.224
5404	0.9708	-0.803	-0.030	-0.028	5203	0.6958	-0.890	-0.030	-0.206
4591.3191	0.6920	-0.877	+0.006	-0.189	5273	0.6959	-0.875	-0.004	-0.211
3299	0.6922	-0.833	-0.046	-0.175	5336	0.6960	-0.894	-0.033	-0.185
3379	0.6924	-0.891	-0.004	-0.171	5408	0.6962	-0.911	-0.030	-0.180
3447	0.6925	-0.868	-0.024	-0.193	5481	0.6963	-0.901	-0.057	-0.179
3518	0.6926	-0.863	-0.041	-0.192	5549	0.6964	-0.890	-0.042	-0.186
3585	0.6928	-0.870	-0.032	-0.188	5616	0.6965	-0.914	-0.018	-0.196
3664	0.6929	-0.864	-0.024	-0.183	5686	0.6967	-0.916	-0.017	-0.198
3735	0.6930	-0.876	-0.039	-0.186	5770	0.6968	-0.918	-0.020	-0.204
3803	0.6932	-0.876	-0.033	-0.185	5890	0.6971	-0.933	-0.003	-0.200
3870	0.6933	-0.876	-0.021	-0.203	5954	0.6972	-0.936	-0.004	-0.207

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
4591.6028	0.6973	-0.916	-0.028	-0.205	4604.5410	0.9379	-0.885	-0.042	-0.061
.6097	0.6974	-0.923	-0.043	-0.207	.5622	0.9363	-0.884	-0.028	-0.086
.6167	0.6976	-0.934	-0.010	-0.220	.5787	0.9386	-0.865	-0.017	-0.063
.6977	0.6238	-0.928	-0.006	-0.211	.5960	0.9390	-0.869	-0.018	-0.063
4602.5133	0.9002	-0.893	-0.049	-0.188	4614.3120	0.1197	-0.935	-0.030	-0.168
.5279	0.9005	-0.896	-0.048	-0.190	.3256	0.1199	-0.937	-0.020	-0.166
.5429	0.9008	-0.902	-0.036	-0.192	4617.2800	0.1749	-0.878	-0.077	-0.144
.5578	0.9011	-0.896	-0.015	-0.217	.2897	0.1751	-0.876	-0.069	-0.164
.5714	0.9013	-0.890	-0.038	-0.206	.2997	0.1752	-0.885	-0.060	-0.164
.5863	0.9016	-0.876	-0.066	-0.204	.3097	0.1754	-0.885	-0.066	-0.172
.6035	0.9019	-0.883	-0.044	-0.188	.3174	0.1756	-0.908	-0.052	-0.163
4603.3419	0.9156	-0.948	-0.059	-0.162	4619.3552	0.2135	-0.918	-0.042	-0.179
.3577	0.9159	-0.940	-0.037	-0.168	.3632	0.2136	-0.934	-0.032	-0.194
.3820	0.9164	-0.945	-0.033	-0.160	.3714	0.2138	-0.911	-0.059	-0.180
.4711	0.9180	-0.918	-0.039	-0.153	.3788	0.2139	-0.923	-0.049	-0.175
.4865	0.9183	-0.923	-0.027	-0.156	.3892	0.2141	-0.945	-0.064	-0.158
.5021	0.9186	-0.923	-0.036	-0.153	.3958	0.2142	-0.933	-0.063	-0.171
.5207	0.9190	-0.921	-0.027	-0.162	.4023	0.2143	-0.933	-0.059	-0.184
4604.3470	0.9343	-0.846	-0.057	-0.089	.4080	0.2145	-0.966	-0.032	-0.153
.3643	0.9347	-0.850	-0.042	-0.090	.4135	0.2146	-0.915	-0.073	-0.176
.3816	0.9350	-0.850	-0.034	-0.086	.4184	0.2146	-0.921	-0.082	-0.187
.4022	0.9354	-0.863	-0.038	-0.082	.4233	0.2147	-0.935	-0.073	-0.166
.4157	0.9356	-0.865	-0.043	-0.073	4625.4046	0.3260	-0.923	-0.059	-0.173
.4325	0.9359	-0.853	-0.041	-0.068	.4204	0.3263	-0.924	-0.055	-0.173
.4463	0.9362	-0.858	-0.034	-0.069	4627.4073	0.3632	-0.899	-0.063	-0.178
.4671	0.9366	-0.877	-0.017	-0.067	.4164	0.3634	-0.906	-0.050	-0.177
.4854	0.9369	-0.864	-0.030	-0.065	.4231	0.3635	-0.904	-0.064	-0.171
.5043	0.9373	-0.873	-0.018	-0.059	.4313	0.3637	-0.900	-0.064	-0.171
.5232	0.9376	-0.865	-0.048	-0.050	.4379	0.3638	-0.901	-0.064	-0.173

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
4627.4455	0.3639	-0.896	-0.074	-0.163	4663.2419	0.0297	-0.771	-0.033	-0.034
4629.3528	0.3994	-0.897	-0.091	-0.179	4692.2492	0.0298	-0.807	-0.019	-0.034
3595	0.3995	-0.909	-0.083	-0.170	4722.2562	0.0300	-0.813	-0.021	-0.027
3657	0.3997	-0.917	-0.074	-0.172	4752.2619	0.0301	-0.809	-0.036	-0.041
3721	0.3998	-0.914	-0.071	-0.172	4782.3070	0.7563	-0.920	-0.062	-0.160
3778	0.3999	-0.896	-0.090	-0.173	4812.3123	0.7564	-0.943	-0.061	-0.163
3837	0.4000	-0.909	-0.085	-0.161	4842.3176	0.7565	-0.936	-0.055	-0.160
3897	0.4001	-0.904	-0.101	-0.168	4872.3215	0.7565	-0.937	-0.055	-0.160
4634.3847	0.4930	-0.903	-0.070	-0.192	4902.2876	0.7745	-0.872	-0.079	-0.163
3917	0.4931	-0.895	-0.072	-0.207	4932.2922	0.7746	-0.860	-0.077	-0.177
3956	0.4932	-0.897	-0.070	-0.182	4962.2951	0.7746	-0.856	-0.062	-0.174
4002	0.4933	-0.894	-0.072	-0.199	4992.3005	0.7747	-0.850	-0.074	-0.179
4635.3116	0.5102	-0.581	-0.052	-0.237	5022.2673	0.0189	-0.648	-0.057	-0.064
3177	0.5103	-0.565	-0.067	-0.240	5052.2737	0.0190	-0.633	-0.072	-0.038
3212	0.5104	-0.572	-0.069	-0.241	5082.2787	0.0191	-0.622	-0.067	-0.063
3263	0.5105	-0.579	-0.079	-0.247	5112.2838	0.0192	-0.637	-0.069	-0.032
4638.4855	0.5693	-0.800	-0.104	-0.186	5142.2893	0.0193	-0.617	-0.073	-0.069
4905	0.5694	-0.827	-0.060	-0.166	5172.3040	0.0196	-0.640	-0.055	-0.058
4950	0.5694	-0.827	-0.090	-0.147	5202.3097	0.0197	-0.650	-0.064	-0.030
4996	0.5695				5232.3150	0.0198	-0.655	-0.053	-0.035
4654.2955	0.8633	-0.870	-0.065	-0.139	5262.3207	0.0199	-0.653	-0.044	-0.053
3040	0.8635	-0.852	-0.070	-0.148	5292.3262	0.0200	-0.648	-0.056	-0.063
3122	0.8636	-0.845	-0.065	-0.156	5322.3374	0.0202	-0.640	-0.050	-0.053
3207	0.8638	-0.845	-0.084	-0.136	5352.3394	0.0203	-0.664	-0.083	-0.033
3271	0.8639	-0.842	-0.072	-0.153	5382.4107	0.0216	-0.647	-0.100	-0.025
4661.3537	0.9946	-0.678			5412.4150	0.0217	-0.651	-0.075	-0.012
3587	0.9947	-0.659	-0.090	-0.037	5442.4194	0.0217	-0.659	-0.084	-0.035
3717	0.9949	-0.675	-0.085	-0.002	5472.4236	0.0218	-0.686	-0.077	-0.022
3758	0.9950	-0.686	-0.056	-0.017	5502.4328	0.0220	-0.669	-0.093	-0.036

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
4985.4431	0.0222	-0.660	-0.070	-0.044	5308.4833	0.0303	-0.765	-0.053	-0.062
.4490	0.0223	-0.658	-0.079	-0.038	.4904	0.0304	-0.768	-0.054	-0.075
.4558	0.0224	-0.686	-0.065	-0.022	.5016	0.0306	-0.767	-0.054	-0.080
.4609	0.0225	-0.683	-0.069	-0.029	.5098	0.0308	-0.759	-0.044	-0.080
.4663	0.0226	-0.692	-0.052	-0.033	.5182	0.0309	-0.765	-0.056	-0.053
.4772	0.0228	-0.689	-0.070	-0.022	.5295	0.0311	-0.771	-0.055	-0.067
.4819	0.0229	-0.679	-0.061	-0.025	.5361	0.0313	-0.757	-0.059	-0.080
.4868	0.0230	-0.694	-0.054	-0.019	.5429	0.0314	-0.758	-0.060	-0.074
.4917	0.0231	-0.689	-0.075	-0.022	.5508	0.0315	-0.745	-0.057	-0.083
.4961	0.0232	-0.687	-0.084	-0.024	.5571	0.0317	-0.718	-0.090	-0.090
.5043	0.0233	-0.678	-0.084	-0.027	.5707	0.0319	-0.729	-0.070	-0.086
.5118	0.0235	-0.691	-0.055	-0.031	.5774	0.0320	-0.748	-0.061	-0.086
.5179	0.0236	-0.711	-0.049	-0.014	.5848	0.0322	-0.762	-0.048	-0.088
.5232	0.0237	-0.727	-0.035	-0.032	.5924	0.0323	-0.767	-0.056	-0.084
.5284	0.0238	-0.714	-0.060	-0.003	.5994	0.0324	-0.768	-0.054	-0.082
.5396	0.0240	-0.710	-0.073	-0.003	.6057	0.0326	-0.768	-0.046	-0.086
.5456	0.0241	-0.709	-0.063	-0.018	.6155	0.0327	-0.751	-0.066	-0.071
.5528	0.0242	-0.693	-0.084	-0.044	.6217	0.0329	-0.749	-0.069	-0.063
.5585	0.0243	-0.717	-0.052	-0.055	.6277	0.0330	-0.754	-0.056	-0.079
.5632	0.0244	-0.693	-0.070	-0.068	.6342	0.0331	-0.765	-0.049	-0.083
.5718	0.0246	-0.689	-0.041	-0.066	.6404	0.0332	-0.771	-0.033	-0.076
.5768	0.0247	-0.689	-0.067	-0.051	.6584	0.0335	-0.772	-0.055	-0.066
.5814	0.0248	-0.719	-0.044	-0.063	.6649	0.0337	-0.789	-0.041	-0.058
.5858	0.0248	-0.701	-0.085	-0.063	.6729	0.0338	-0.784	-0.049	-0.056
5308.4426	0.0295	-0.750	-0.070	-0.070	.6801	0.0339	-0.789	-0.062	-0.067
.4510	0.0297	-0.757	-0.052	-0.085	.6884	0.0341	-0.771	-0.038	-0.095
.4579	0.0298	-0.759	-0.057	-0.066	5309.3959	0.0473	-0.755	-0.055	-0.053
.4647	0.0299	-0.758	-0.047	-0.075	.4018	0.0474	-0.760	-0.057	-0.051
.4738	0.0301	-0.764	-0.048	-0.069	.4080	0.0475	-0.759	-0.054	-0.071

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$	$\Delta(U-B)$
5309.4149	0.0476	-0.765	-0.055	-0.058	5309.6257	0.0515	-0.832	-0.034	-0.072
.44214	0.0477	-0.775	-0.048	-0.051	.6349	0.0517	-0.830	-0.032	-0.068
.4298	0.0479	-0.786	-0.041	-0.057	.6434	0.0519	-0.823	-0.033	-0.073
.4361	0.0480	-0.791	-0.025	-0.067	.6508	0.0520	-0.834	-0.037	-0.060
.4424	0.0481	-0.814	+0.006	-0.058	.6579	0.0521	-0.834	-0.061	-0.079
.4487	0.0482	-0.774	-0.052	-0.070	.6655	0.0523	-0.843	-0.030	-0.068
.4539	0.0483	-0.770	-0.057	-0.071	.6721	0.0524	-0.817	-0.043	-0.066
.4604	0.0485	-0.820	-0.008	-0.054	.6805	0.0526	-0.809	-0.035	-0.077
.4688	0.0486	-0.807	-0.004	-0.060	.6864	0.0527	-0.809	-0.049	-0.077
.4766	0.0488	-0.779	+0.006	-0.085	.6928	0.0528	-0.817	-0.075	-0.065
.4841	0.0489	-0.781	-0.039	-0.066	5782.3232	0.8430	-0.873	-0.007	-0.211
.4910	0.0490	-0.795	+0.002	-0.070	.3330	0.8432	-0.904	+0.016	-0.227
.4989	0.0492	-0.780	+0.013	-0.100	.3421	0.8434	-0.879	-0.004	-0.213
.5060	0.0493	-0.815	0.000	-0.054	.3665	0.8438	-0.869	-0.019	-0.209
.5168	0.0495	-0.864	+0.015	-0.038	.3757	0.8440	-0.876	+0.010	-0.251
.5238	0.0496	-0.796	-0.053	-0.055	.3865	0.8442	-0.892	-0.002	-0.233
.5300	0.0498	-0.783	-0.028	-0.075					
.5361	0.0499	-0.767	-0.050	-0.080					
.5424	0.0500	-0.791	-0.033	-0.075	4590.7787	0.6820	-0.869		
.5525	0.0502	-0.793	-0.056	-0.048	.7898	0.6822	-0.887		
.5607	0.0503	-0.815	-0.011	-0.049	.8162	0.6827	-0.873		
.5686	0.0505	-0.803	-0.029	-0.073	4634.6603	0.4981	-0.820		
.5757	0.0506	-0.801	-0.053	-0.086	.6943	0.4988	-0.836		
.5846	0.0508	-0.806	-0.048	-0.069	.7130	0.4991	-0.799		
.5933	0.0509	-0.812	-0.022	-0.074	4635.6463	0.5165	-0.730		
.5994	0.0510	-0.815	-0.045	-0.085	.6754	0.5170	-0.695		
.6061	0.0512	-0.807	-0.045	-0.082	.6865	0.5172	-0.680		
.6128	0.0513	-0.818	-0.048	-0.075	4662.5877	0.0175	-0.645		
.6191	0.0514	-0.811	-0.051	-0.088	.6134	0.0180	-0.670		

Waterloo

Table 1 (continued)

J.D. hel 2440000+	Phase	ΔV	J.D. hel 2440000+	Phase	ΔV	$\Delta(B-V)$
4662.6280	0.0183	-0.686	4701.5801	0.7427	-0.916	
.6481	0.0187	-0.680	.5912	0.7429	-0.925	
.6752	0.0192	-0.678	4702.5640	0.7610	-0.945	
4667.5623	0.1101	-0.930	.5758	0.7613	-0.930	
.5804	0.1104	-0.922	.5862	0.7614	-0.912	
.5984	0.1107	-0.915	.5932	0.7616	-0.930	
.6151	0.1110	-0.913	4704.5457	0.7979	-0.932	
.6429	0.1116	-0.930	.5575	0.7981	-0.945	
4670.6273	0.1671	-0.980	.5652	0.7983	-0.922	
.6398	0.1673	-0.983	.5714	0.7984	-0.947	
.6606	0.1677	-0.990	.5791	0.7985	-0.955	
4681.5820	0.3708	-0.880	4912.8985	0.6730	-0.947	-0.012
4682.5985	0.3897	-0.865	4921.8770	0.8399	-0.881	-0.018
4689.5888	0.5197	-0.808	.9041	0.8405	-0.891	-0.014
.6096	0.5201	-0.815	4933.9042	0.0636	-0.880	+0.018
.6214	0.5203	-0.822	4998.6694	0.2682	-0.860	-0.006
4691.6039	0.5572	-0.915	5011.6152	0.5090	-0.599	+0.014
.6268	0.5576	-0.908	5026.5774	0.7872	-0.914	-0.035
4701.5641	0.7424	-0.910	5027.5780	0.8058	-0.897	-0.041

luminosity class from dwarf to bright supergiant and back, maintaining however its photometric spectral subclass.

The photometric light curve given in Fig. 1 challenges the model of V 505 Mon as a detached system (Stagni et al., 1982), because it shows some indications of ongoing mass transfer in a semidetached binary configuration. The asymmetries described in previous section can be well explained by the existence of accretion disk surrounding the hot component, gaseous stream from cooler component and hot impact region on the accretion disk. The orbital period is long, so the presence of the disk in a semidetached system is anticipated. H α emission detected by Chochol and Kučera (1981) and Stagni et al. (1982) may arise in the accretion disk. The observed short-term changes may be due to orbital motion of inhomogenities in accretion disk, as was proposed by Bath (1977). If this is the case, it is not necessary to expect strict periodicity.

The further spectroscopic study of the system is highly desirable.

ACKNOWLEDGEMENTS

The authors wish to express their thanks to Dr. P. Harmanec for valuable comments and suggestions and the staff members of the stellar department of the Skalnaté Pleso Observatory for practical help.

REFERENCES

- Bath, G. T.: 1977, *Monthly Notices Roy. Astron. Soc.* 178, 203.
Chochol, D.: 1983, in *Stellar Spectra and their interpretation*, (Eds.: I. Hubený and B. Onderlička), *Publ. Astron. Inst. Czechosl. Academy of Sci.*, No. 57, p. 30.
Chochol, D., Kučera, A.: 1981, *Inform. Bull. Var. Stars*, No. 1998.
De Gréve, J. P., Doom, G., Hellings, P., Packet, W., Stagni, R., Mammano, A.: 1983, *Astrophys. Space Sci.* 95, 215.
Harmanec, P.: 1983 a, in *Advances in Photoelectric Photometry*, Vol. 1, (Eds.: R. C. Wolpert, R. M. Genet), Fairborn Observatory, p. 42.
Harmanec, P.: 1983 b, in *Rapid Variability of Early-type Stars*, (Eds.: P. Harmanec and K. Pavlovský), Hvar Obs. Bull. 7, 55.
Hoag, A. A., Smith, E. P.: 1959, *Publ. Astron. Soc. Pacific* 71, 32.
Stagni, R., Margoni, R., Mammano, A.: 1982, *Astrophys. Space Sci.* 88, 115.
Stellingwerf, R. F.: 1978, *Astrophys. J.* 224, 953.
Turner, D. G.: 1976, *Astrophys. J.* 210, 65.
Wachmann, A. A.: 1966, *Astron. Abhandl. Hamburger Sternw.* 7, 341.